

2D Vacuum transitions and their holographic interpretation

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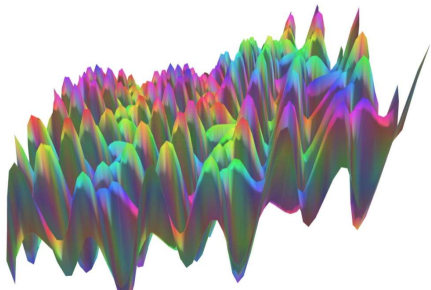
supervised by Professor Fernando Quevedo



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Outline:

1. Motivations
2. methods: Euclidean (CDL, BT), and Hamiltonian (FMP)
3. **Holographic** interpretation



Motivations

1. populating the String Theory Landscape
2. promising for understanding QG (analogy with BHs)
3. testing the consistency of the 3 formalisms
4. similar unitarity issues as in the information loss paradox
5. analytic calculation in 2D.

1) CDL₂ from JT-gravity using Almheiri-Polchinski

$$S_{2D} = \int d^2x \sqrt{-g} [\phi^2 \mathcal{R} + \lambda(\phi')^2 - U(\phi)]$$

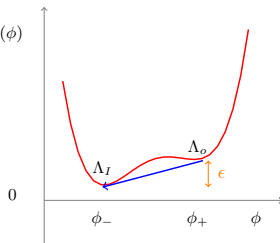
$$ds^2 = e^{2\omega} (dt^2 + dz^2) = \rho^2 (dr^2 + r^2 d\theta^2) \quad V(\phi)$$

$$S_E = 2\pi \int_0^{\bar{r}} dr \left[2\phi^2 (r\omega')' + r U_o e^{2\omega} \right]$$

1. the vacua $(\phi^2)' = 0$, $\partial_{\phi^2} U_o \stackrel{\text{def}}{=} \Lambda = \text{const.}$

2. the wall $(\phi^2)' \neq 0 \Rightarrow B_{\text{wall}}$

constraint: $(\phi^2)' = c_1 r e^{2\omega}$



$$\Gamma \sim \exp(iS_o) \stackrel{\text{def.}}{=} \exp(-B) \quad , \quad B \stackrel{\text{def.}}{=} S_E|_{\text{inst}} - S_E|_{\text{bckgr}}$$

$$\text{e.g.} \quad B_{\text{tot}}^{\text{AdS}_2 \rightarrow \text{AdS}_2} = 4\pi \left(\frac{\phi_-^2}{\phi_+^2} + 2 \right) (\phi_-^2 - \phi_+^2)$$

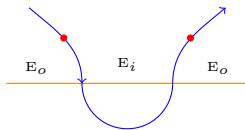
$$B_{\text{tot}} \stackrel{\text{def.}}{=} B_{\text{in}} + B_{\text{wall}}$$

2) BT: brane nucleation ($D=d+1$), fits well with the Landscape: no V

In 2D : 1. Membrane \equiv particle/anti-particle pair.
2. no magnetic field.

Effectively

$$\Lambda_{o,i} \stackrel{\text{def.}}{=} \lambda + \frac{1}{2} k E_{o,i}^2$$



$$B_{tot}^{type-1} \Big|_{extr} = \frac{4\pi}{k} \ln \left| \frac{r_i}{r_o} \right|$$

$$r_{o,i} \stackrel{\text{def.}}{=} \frac{2}{\Lambda_{o,i} \bar{\rho}} \left[1 - \sigma_{o,i} \sqrt{1 - \Lambda_{o,i} \bar{\rho}^2} \right]$$

$$\sigma_{o,i} \stackrel{\text{def.}}{=} \text{sgn} \left[\mathcal{E} e E_{on} \mp \frac{1}{4} k m^2 \right] \quad , \quad \bar{\rho} \stackrel{\text{def.}}{=} \frac{m}{\sqrt{\Lambda_i m^2 + (\mathcal{E} e E_{on} + \frac{1}{4} k m^2)^2}}$$

Some preliminary entropic arguments

$$\frac{\Gamma_{t \rightarrow f}}{\Gamma_{f \rightarrow t}} = \exp \left[-\frac{3}{8} M_P^4 \left(\frac{1}{V_t} - \frac{1}{V_f} \right) \right] \quad [\text{Lee-Weinberg, 1987}]$$

1. Not for $V_t < 0$ due to topology change.

2. $\Rightarrow \frac{\Gamma_{o \rightarrow I}}{\Gamma_{I \rightarrow o}} = e^{\frac{\eta\pi}{2G} \left(\frac{1}{H_o^2} - \frac{1}{H_I^2} \right)} = e^{\eta(S_o - S_I)} \xrightarrow{H_o \rightarrow 0} 0$, since $S_{bckgr} \rightarrow \infty$

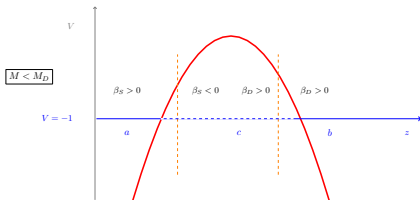
3. $S_{\text{Sch} \rightarrow \text{dS}} = \frac{\eta\pi}{2G} [S_{\text{bounce}} - S_{\text{bckgr}}] = \frac{\eta\pi}{2G} [S_{\text{bounce}} - 4G^2 M^2]$

Sch₄ → dS₄ [FGG]

If $M < M_D$

$\lim_{M_S \rightarrow 0} \Rightarrow$ brane nucleation

from vacuum fluctuations



1. Avoid the **initial singularity problem** by quantum tunnelling, but,
2. the instanton is singular. A **pseudo-manifold** [motivating FMP]

3) The Hamiltonian formulation of VTs [FMP]

1. Lorentzian

2. in 2D using JT.

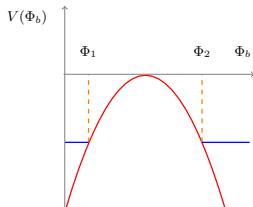
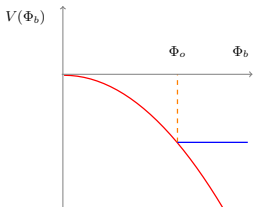
$$g_{\mu\nu} = \begin{pmatrix} -(N^t)^2 + (N^z)^2 & LN^z \\ LN^z & L^2 \end{pmatrix}$$

$$P(A \rightarrow B) \stackrel{\text{def.}}{=} \frac{|\Psi(\text{nothing} \rightarrow A/B)|^2}{|\Psi(\text{nothing} \rightarrow A)|^2}$$

$$\begin{aligned} \mathcal{L}_{2D} &= \sqrt{-g} \phi [\mathcal{R} - 2\Lambda_+ \Theta(z - \hat{z}) - 2\Lambda_- \Theta(\hat{z} - z)] + \\ &\quad - \delta(z - \hat{z}) \sigma \sqrt{(N^t)^2 + L^2 (N^z + \dot{\hat{z}})^2} + \\ &\quad + \sqrt{-g} [\phi_o \mathcal{R} - \mathcal{B}_+ \Theta(z - \hat{z}) - \mathcal{B}_- \Theta(\hat{z} - z)] \\ &= \mathcal{L}_{BTs} + \pi_L \dot{L} + \pi_\phi \dot{\phi} - N^t \mathcal{H}_t - N^z \mathcal{H}_z \end{aligned}$$

$$\mathcal{H}_z = 0 \Big|_{z \neq \hat{z}} \Rightarrow \boxed{\pi'_L = \phi' \frac{\pi_\phi}{L}} \quad \text{If } N^t = 1, \ N^z = 0, \ \sigma = \kappa \phi, \ \Phi_b \stackrel{\text{def.}}{=} \frac{\phi_b}{\sqrt{|C|}}$$

$$\left\{ \begin{array}{l} - \int_{\hat{z}-\epsilon}^{\hat{z}+\epsilon} dz \mathcal{H}_z = 0 \\ - \int_{\hat{z}-\epsilon}^{\hat{z}+\epsilon} dz \mathcal{H}_t = 0 \end{array} \right. \Rightarrow \left\{ \begin{array}{l} \left. \frac{\phi'}{L} \right|_{\hat{z}_-}^{\hat{z}_+} = -\frac{\sigma}{2} \\ \left. \pi_L \right|_{\hat{z}_-}^{\hat{z}_+} = 0 \end{array} \right. \Rightarrow \boxed{\dot{\Phi}_b^2 + V_{eff} = -1}$$



$$ds^2 = \frac{\left(\phi \pm \frac{\mathcal{B}}{2\Lambda_{\pm}}\right)^2 - 4a^2}{4a^2} \left(-dt^2 + \frac{4a^2}{b^2} \frac{d\phi^2}{\left(\left(\phi \pm \frac{\mathcal{B}}{2\Lambda_{\pm}}\right)^2 - 4a^2\right)^2} \right)$$

Main results

$$S_{tot}^{\text{AdS}_2 \rightarrow \text{AdS}_2} \Big|_{\mathcal{B}=0} = \frac{\eta}{G} \left[\mu_- \ln \left| \frac{\mu_- A_1 + 1}{\mu_- A_1 - 1} \right| - \mu_+ \ln \left| \frac{\mu_+ A_2 + 1}{\mu_+ A_2 - 1} \right| \right] \equiv S_{brane}$$

$$S_{tot}^{\text{dS}_2 \rightarrow \text{dS}_2} \Big|_{\mathcal{B}=0} = \frac{2\eta}{G} \pi (\mu_+ - \mu_-) + S_{brane} \xrightarrow{\mu_- \rightarrow +\infty} +\infty$$

$$S_{tot}^{\text{Mink}_2 \rightarrow (\text{A})\text{dS}_2} \Big|_{\mathcal{B} \neq 0} = \frac{2\eta}{G} \left[\phi_f - \phi_{h,2}^- + \rho_2 - \rho_1 + \right. \\ \left. + L_3 \left[\tan(h) \left(\frac{\rho_2}{L_3} \right) \ln \left| A_2 \sin(h) \left(\frac{\rho_2}{L_3} \right) \right| - \tan(h) \left(\frac{\rho_1}{L_3} \right) \ln \left| A_2 \sin(h) \left(\frac{\rho_1}{L_3} \right) \right| \right] \right]$$

Notation: $\mu_{\pm} \stackrel{\text{def.}}{=} \sqrt{-\frac{\mathcal{C}}{\Lambda_{\pm}}}$, $A_{1,2} \stackrel{\text{def.}}{=} \frac{\Lambda_+ - \Lambda_- \pm \frac{\kappa^2}{4}}{\kappa}$, $L_3^2 \stackrel{\text{def.}}{=} \frac{\frac{\mathcal{B}^2}{4\Lambda_-} - \mathcal{C}}{\Lambda_-}$

$\lim_{\Lambda \rightarrow 0} \phi_h \Big|_{\mathcal{B} \neq 0} \stackrel{\text{def.}}{=} \phi_f = \text{const.}$, $y \stackrel{\text{def.}}{=} \tanh \left(\frac{\rho}{L_3} \right) \stackrel{\text{def.}}{=} L_3^{-1} \left(\frac{\mathcal{B}}{2\Lambda_-} + \phi \right)$

BT as $\text{AdS}_2/\text{CFT}_1 \subset \text{AdS}_3/\text{CFT}_2$

2 types of Einstein -Maxwell dilaton gravity [Hartman, Strominger, 2008; Alhishahia, , 2008] :

1. **type-I** $S = \frac{1}{16\pi G} \int dx^2 \sqrt{-g} \left(e^\phi \left(\mathcal{R} + \frac{2}{\ell^2} \right) + F^2 \right)$

2. **type-II** $S = \frac{1}{16\pi G} \int dx^2 \sqrt{-g} e^\phi \left(\left(\mathcal{R} + \frac{2}{\ell^2} \right) + e^{2\phi} F^2 \right)$

Both admit an AdS_2 solution dual to a CFT_1 , but only type-II can be obtained by dimensional reduction of $\text{AdS}_3/\text{CFT}_2$ where $c \stackrel{\text{def}}{=} \frac{3R_3}{2G}$.

BT is a type-II theory.

New perspective:

a) looking for a possible entropic understanding of B_{tot} or S_{tot} going beyond detailed balance using holography;

b) possible island interpretation.

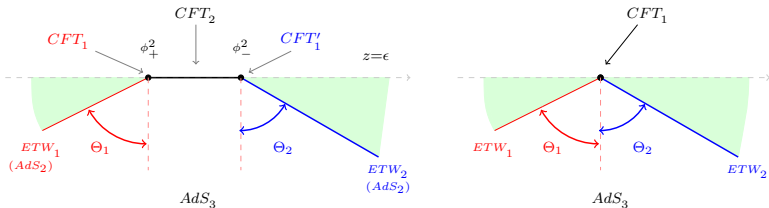
Holographic interpretation: examples

$$\begin{aligned}
 \text{A)} \quad B_{tot}^{\text{AdS}_2 \rightarrow \text{AdS}_2} &= 4\pi \left(\frac{\phi_-^2}{\phi_+^2} + 2 \right) (\phi_-^2 - \phi_+^2) = \\
 &\stackrel{\text{def.}}{=} 4\pi S_{\pm}^2 = 4\pi S_+ S_- = \\
 &= 4\pi c_{hol}^{CFT_2} M_{BH}
 \end{aligned}$$

$$S_{\pm} \text{ generalising } S_{\text{Cardy}} = 2\pi \sqrt{\frac{c}{6} \left(\Delta - \frac{c}{24} \right)}$$

$$c_{hol}^{CFT_2} = \frac{\partial(S_+ S_-)}{\partial Q} = \left(\frac{\phi_-^2}{\phi_+^2} + 2 \right) \phi_+^2, \quad M_{BH} = \frac{\phi_-^2}{\phi_+^2} - 1$$

1. the newly-nucleated spacetime is a black hole in the background:
relative entropy
2. the spacetimes are extremal and the RG-flow is the CFT_2 with c_{BCFT_2}
3. $\lim_{\Lambda_{\pm} \rightarrow 0} B_{tot} = \infty$ in agreement with the c-theorem.
4. $\phi_-^2 \equiv \phi_+^2 \Rightarrow B_{tot} = 0$ all the information is in the CFT_2 (see next).



after compactification, it is consistent with [wedge holography](#):
 $AdS_2/CFT_1 \subset AdS_3/CFT_2$ just as BT:

$$B) \left. B_{type-1}^{AdS_2 \rightarrow AdS_2} \right|_{extr} = 2\pi \left[S_{bdy}^- - S_{bdy}^+ - \ln \left| \frac{\Lambda_+}{\Lambda_-} \right| \right] = 2\pi S_{ICFT}$$

with $S_{bdy} \stackrel{def.}{=} \ln \left| \frac{1+T}{1-T} \right|$, $T \stackrel{def.}{=} \sqrt{1 - \Lambda_{o,l} \bar{\rho}^2}$

Key point: T is uniquely determined by ϵ .

Relation to $T\bar{T}$ \Rightarrow Transitions are **local** and the **CFT₂** lives at $z = \epsilon$.
 From FMP

$$S_{tot}^{(A)dS \rightarrow (A)dS} = 2\pi\eta \left[S_{T\bar{T}}^- - S_{T\bar{T}}^+ \right]_{\text{univ.}} = -2\pi\eta F_{\partial} = 2\pi\eta S_{ICFT}$$

$$-F_{\partial} \stackrel{\text{def.}}{=} S_{\partial}(R)_{\text{univ.}} \stackrel{\text{def.}}{=} \lim_{\epsilon \rightarrow 0} \left[\ln \mathcal{Z} \left[HS_R^D \right] - \frac{1}{2} \ln \mathcal{Z} \left[S_R^D \right] \right]$$

$$S_{T\bar{T}} \stackrel{\text{def.}}{=} \left(1 - \frac{r}{2} \frac{\partial}{\partial r} \right) \ln \mathcal{Z}_{S^2}^{(A)dS, T\bar{T}} =$$

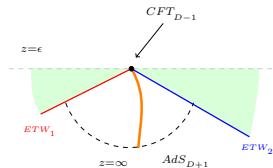
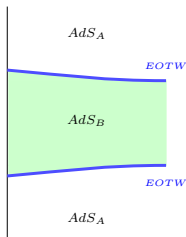
$$= \begin{cases} \pi \frac{\epsilon}{3} \sinh^{-1} \left(\sqrt{\frac{12}{c\lambda}} r \right), & [\text{H. Verlinde et al.}] \\ \frac{\epsilon}{6} + \frac{\epsilon}{3} \ln \left| \sqrt{\frac{\lambda c}{12}} \frac{1}{r} \right| + \frac{\epsilon}{3} \tan^{-1} \left(\frac{1}{\sqrt{\frac{\lambda c}{12r^2} - 1}} \right) + \text{const.}, & [\text{Silverstein et al.}] \end{cases}$$

$$\text{with } \mu \stackrel{\text{def.}}{=} \sqrt{\frac{\lambda c}{12}}, \quad \phi_o \stackrel{\text{def.}}{=} r, \quad \lambda \stackrel{\text{def.}}{=} 2\sqrt{-\frac{c}{\Lambda}}$$

The emergence of the island [equivalent to Van Raamsdonk et al.]

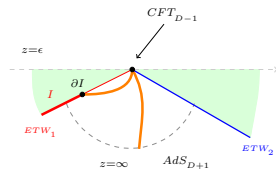
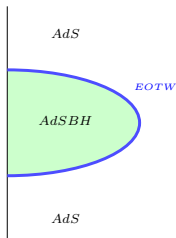
$$\text{AdS}_2/\text{CFT}_1 \subset \text{AdS}_3/\text{CFT}_2$$

$$S_{\text{tot}}^{\text{AdS}_2 \rightarrow \text{AdS}_2} \Big|_{\mathcal{B}=0} \equiv$$



$$\text{AdS}_2/\text{CFT}_1 \not\subset \text{AdS}_3/\text{CFT}_2$$

$$S_{\text{tot}}^{\text{AdS}_2 \rightarrow \text{AdSBH}_2} \Big|_{\mathcal{B} \neq 0} \equiv$$

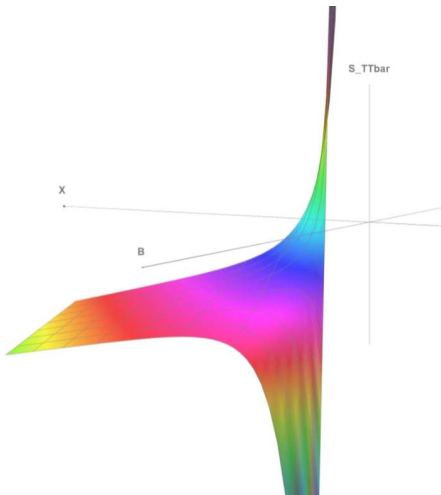


$$S_{brane} = S_{EE}^{(2)} \Big|_{\mathcal{B}=0} + \delta S_{EE}^s + \delta S_{EE}^t + \mathcal{O}(\mathcal{B}^2) \quad , \quad \phi_2 \stackrel{def.}{=} -\frac{\mathcal{B}}{|\mathcal{C}|} \phi_o^2 - \phi_1$$

Phase transition beyond a certain value of \mathcal{B} .

$$S_{brane} = \eta \left(S_{EE,-}^{T\bar{T}} \pm S_{EE,+}^{T\bar{T}} \right)$$

$$\begin{aligned} S_{EE}^{1,2} &= S_{EE}^{(2)} \Big|_{\mathcal{B}=0}^{1,2} + \delta S_{EE}^s \Big|_{\mathcal{B}=0}^{1,2} \\ &\sim -\frac{8c_-}{3} \ln \left| \frac{1}{\epsilon} \sinh \left(\frac{2\pi r_-}{\beta_-} \right) \right| + \\ &\quad - \left(1 - \frac{\phi_o^2}{\mu_-^2} \right) \frac{\mathcal{B}c_-^2}{9\Lambda_-} \coth^2 \left(\frac{2\pi r_-}{\beta_-} \right) + \\ &\quad + \frac{\frac{c_-}{3}}{\sinh^2 \left(\frac{2\pi r_-}{\beta_-} \right)} \end{aligned}$$



	AdS ₂	dS ₂	Mink ₂
	↑	↑	↑
AdS ₂ →	CDL BT FMP	CDL FMP	CDL FMP
dS ₂ →	CDL BT FMP	CDL BT FMP	CDL BT FMP
Mink ₂ →	CDL BT FMP	CDL FMP	

Main results:

1. **Agreement** in between the 3 formalisms in most of the cases, thanks to their **holographic interpretation**
2. **The flat limit** of 2D vacuum transitions requires **a black hole**
3. black holes play a similar role as **$T\bar{T}$** -deformations in a CFT_2
4. the BH **mass** is related to ϵ_{UV} of a CFT_2
5. the behaviour of the total action is similar to the difference of 2 **generalised entropies** [Maldacena et al.]

$$S_{tot}^{(A)dS \rightarrow (A)dS} = 2\pi\eta \left[\frac{\phi_{h,+}}{G} - \frac{\phi_{h,-}}{G} + S_{EE,-}^{(A)dS, T\bar{T}} - S_{EE,+}^{(A)dS, T\bar{T}} \right]$$

Future directions & Work in progress:

1. extension to SUGRA solutions
2. ...